




## SOLAR WIND



Continuous radial outflow of gas.
Consists of charged particles that can carry magnetic structures.

Shapes the IMF.

## INTERPLANETARY MAGNETIC FIELD + PLASMA

## PARKER SPIRAL

The magnetic field stays connected to the Sun. As the Sun rotates, the magnetic field gets bended. Nevertheless, the plasma goes radial.


## SOLAR WIND IS RADIAL \& IMF BENDS?



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## SOLAR WINDVARIATIONS

Non-eruptive -'̀', Eruptive
No B reconnection C reconnection

Coronal Hole


Coronal Mass Ejection


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## B OF CORONAL HOLE



WHAT HAPPENS WHEN FAST CATCHES SLOW?

## Top View


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Interaction Region Rotates together with the Sun

## Coronal Hole

- Co-rotating structure
- Radial!
- No extra mass-flux



## Coronal Hole




## RECONNECTION

The magnetic field carried by the solar wind can couple with the magnetic field of Earth. This coupling is stronger when the solar wind magnetic field is opposite to the magnetic field of Earth.

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# COORDINATE SYSTEM 


$+Z$ is perpendicular to the Ecliptic Plane.


## IMF POLARITY



## EXTREME FAST



## EXTREME FAST



## EXTREME SLOW



Central Meridian


# Beacon COR2 CME speed calculator 

https://wwwbis.sidc.be/cactus/


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