

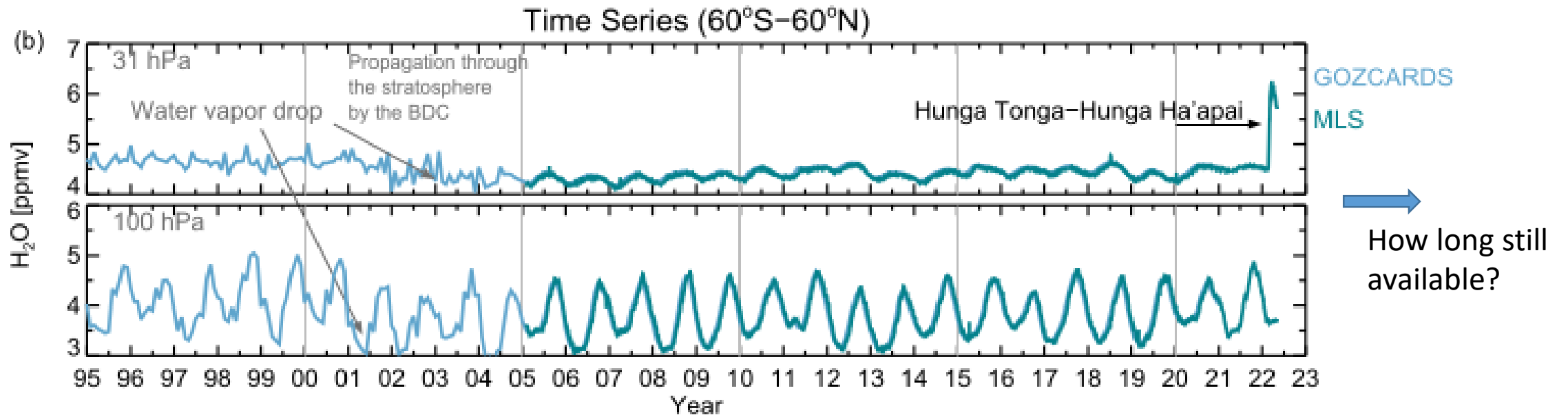
# Investigating possible contributions of IASI/IASI-NG for bridging the upcoming gap of limb sounding observation capabilities

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- Motivation
- Stratospheric WV from thermal nadir observations: problem, idea, possibility, ...
- Performance test
- Some first results (ongoing work...)
- Summary

# Motivation

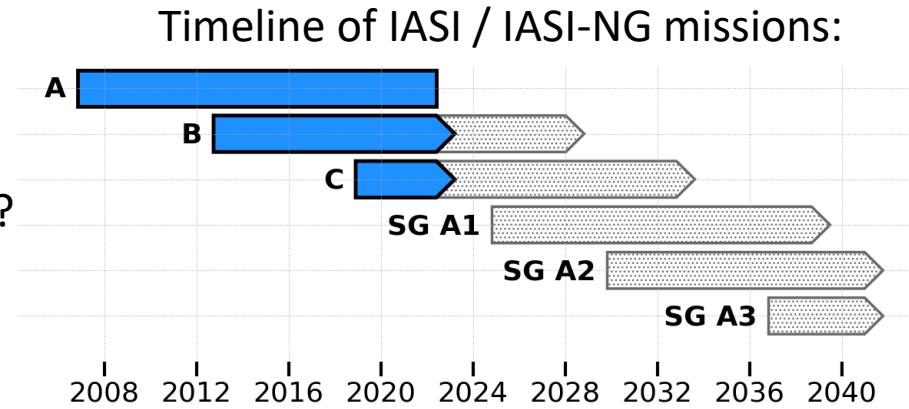
Satellite instruments that measure the atmosphere in limb-viewing geometry are very successful in observing vertical profiles in and above the UTLS (upper troposphere/lower stratosphere), thus providing a unique data pool for investigating the impact of the UTLS on weather and climate. Unfortunately, missions like MIPAS, ACE-FTS, or Aura/MLS have already ended or are expected to end in several years before respective next-generation missions will be operative (e.g. the current ESA Earth Explorer 11 candidate CAIRT). Here we discuss to what extent IASI/IASI-NG can bridge this upcoming observational gap



Millán et al. (2022)

# Motivation + Problem

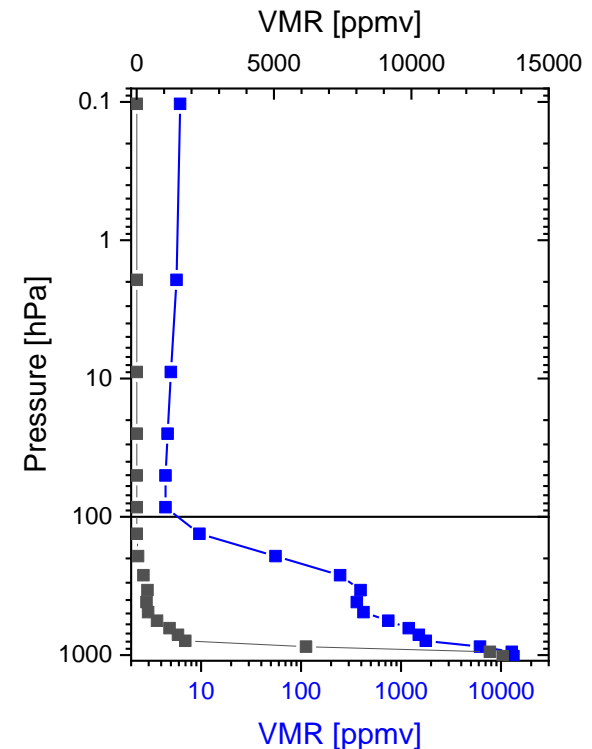
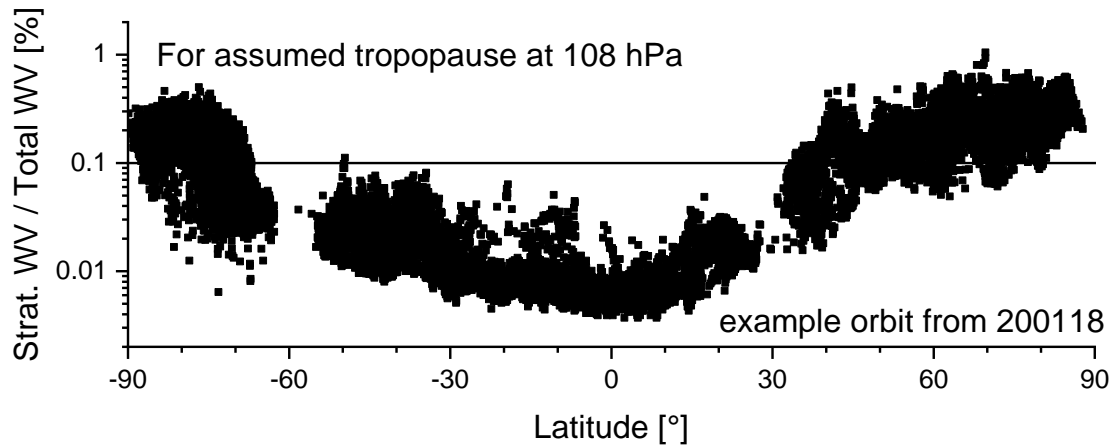
Can thermal nadir observations help to bridge the upcoming gap?  
IASI and IASI-NG is/will be available for many decades.



## Problem:

Nadir viewing geometry, only very small amount of water is in the stratosphere.

Generally more than 99.9% is in the troposphere.



# Approach

... to minimize interferences of variations of tropospheric WV on the stratospheric WV retrieval:

- Avoid an independent temperature and WV retrieval in the troposphere (there is a physical relation between both). Instead make an optimal estimation of RH.
- In the stratosphere humidity is clearly below saturation and RH is not important: independent retrieval of temperature and WV

(1) Retrieval WV on log-scale

(2) According to WMO (Guide to Meteorological Instruments and Methods of Observation,

Appendix 4B, WMO-No. 8 (CIMO Guide), Geneva 2008):  $x_{H_2O}^{sat} = \frac{6.112 e^{\left(\frac{17.62 T}{243.12+T}\right)}}{p}$

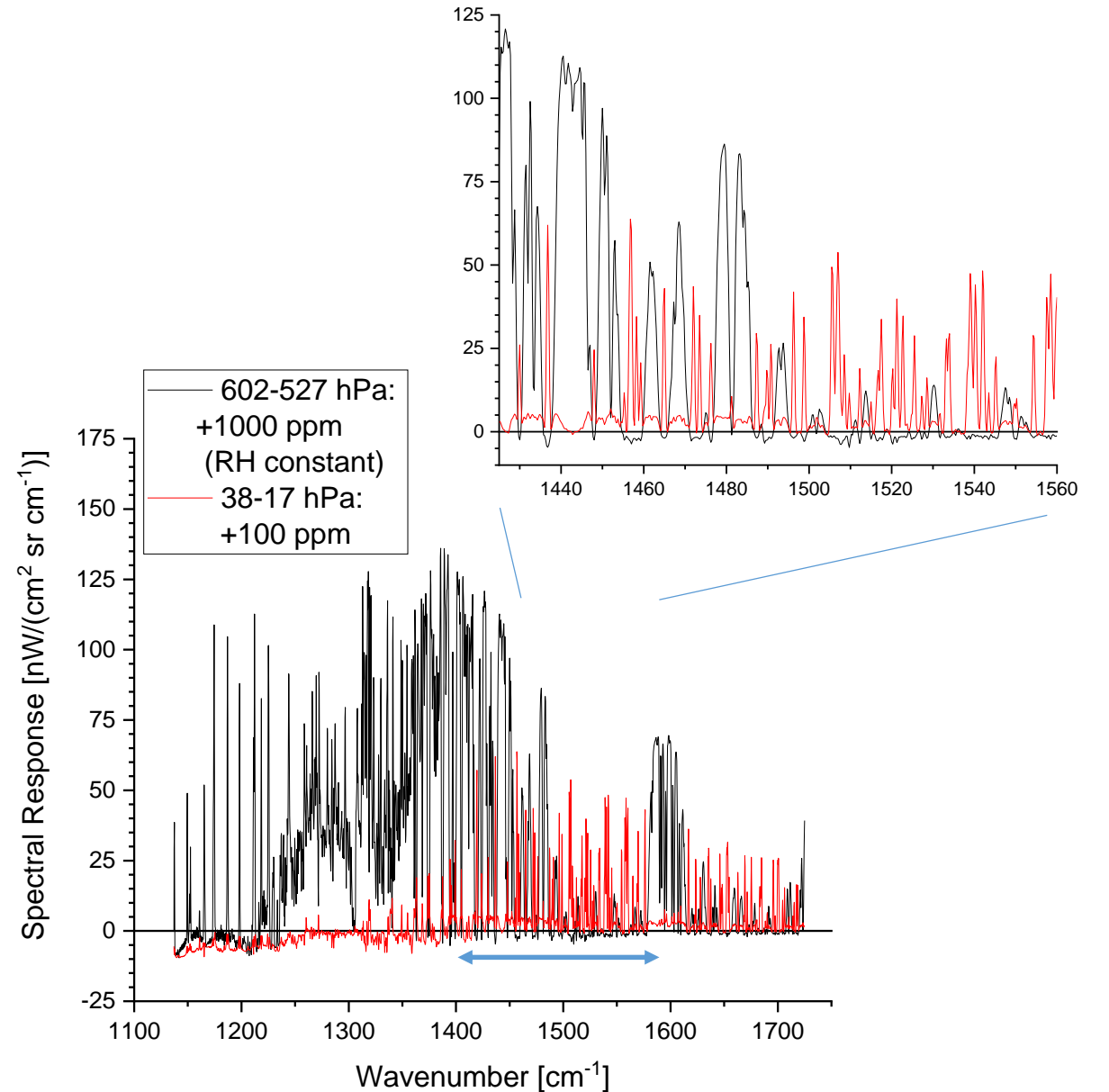
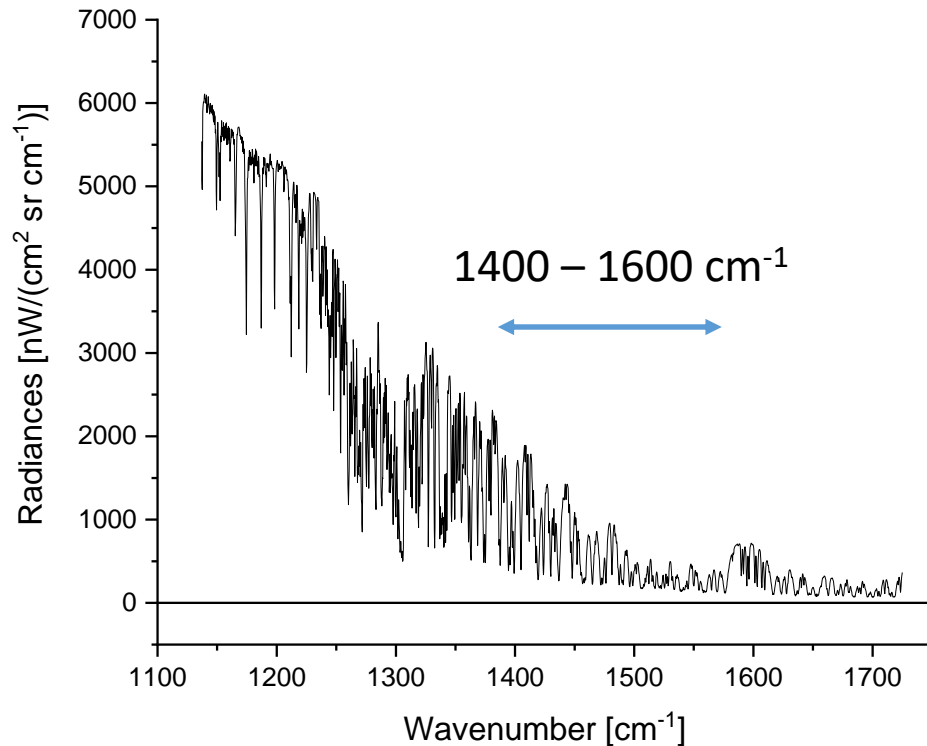
$$\rightarrow \frac{\Delta \ln(x_{H_2O}^{sat})}{\Delta T} = \frac{17.62 * 243.12}{(243.12+T)^2}$$

(3) Cross-constraint between  $\ln(x_{H_2O}^{sat})$  and T is a constraint on RH

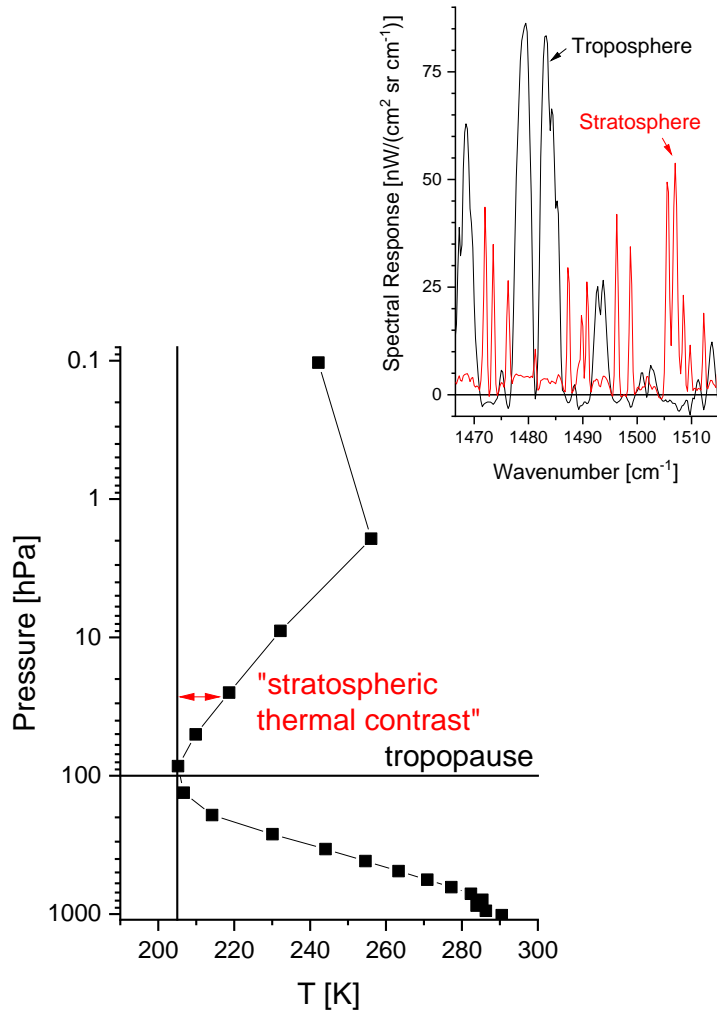
# Spectroscopic responses

Stratospheric WV variations are visible in the spectral region from 1400 to 1600  $\text{cm}^{-1}$

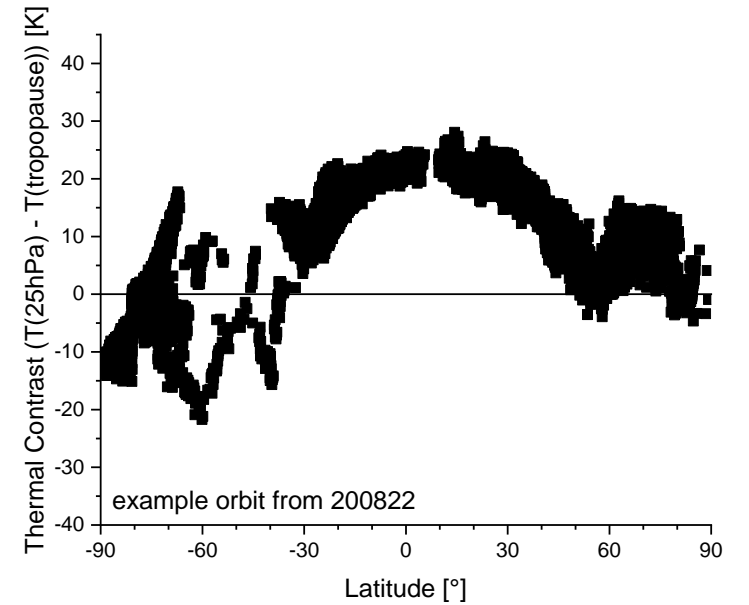
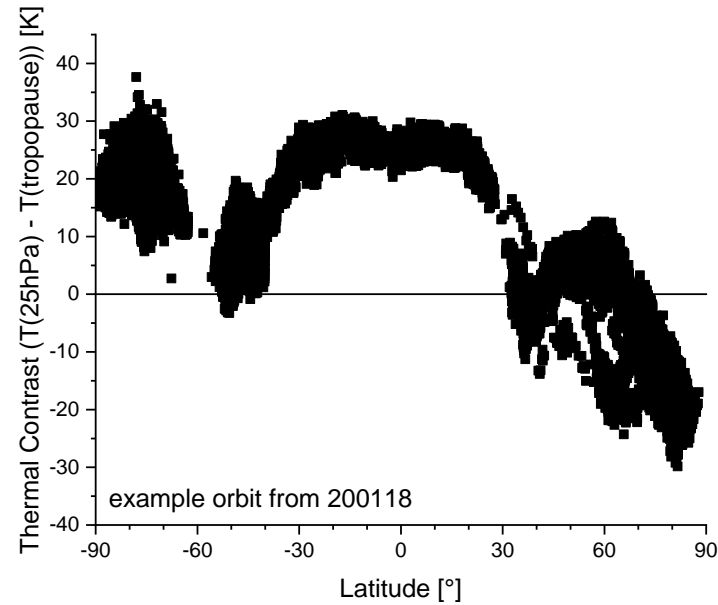
Example for 220112, 32°S:



# Spectroscopic responses

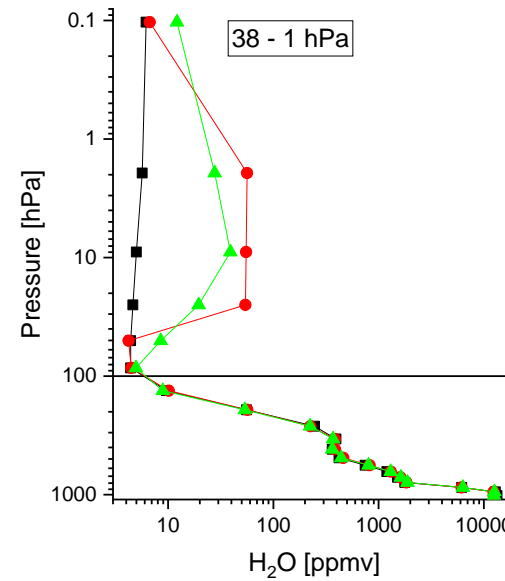
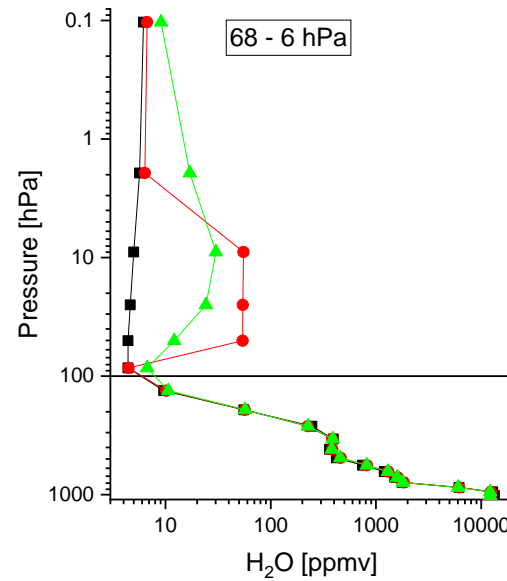
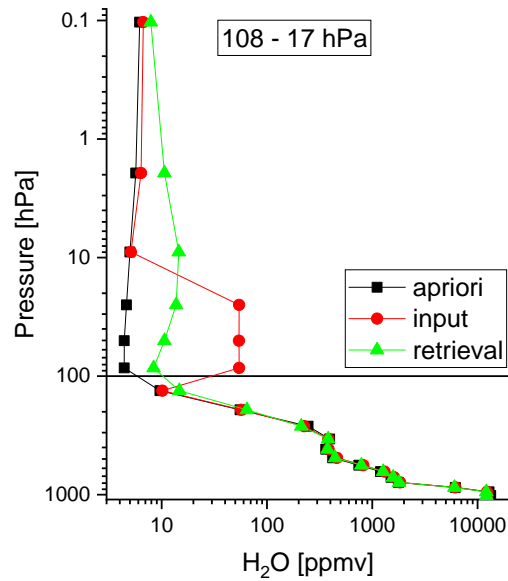


“Stratospheric thermal contrast” is important.  
→ we expect the best stratospheric WV signals in thermal nadir radiances for lower latitudes.



# Performance test

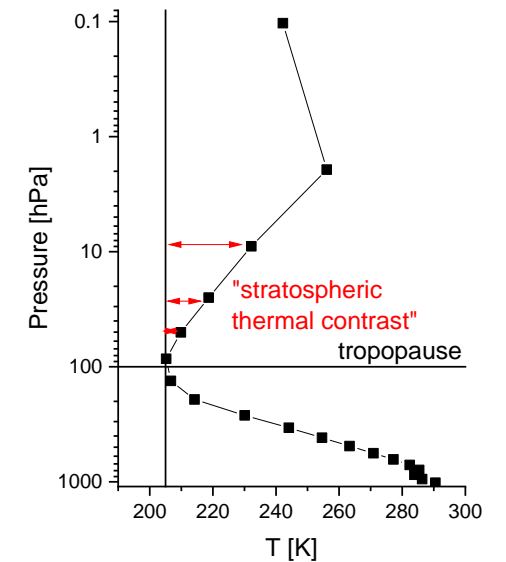
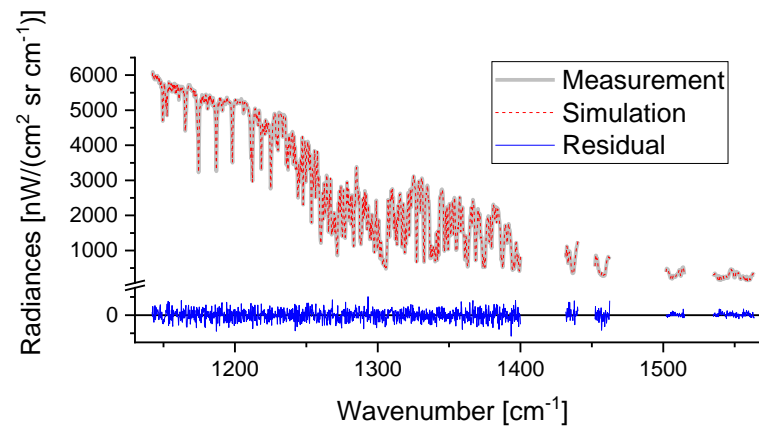
We simulate spectra with typical noise and different stratospheric WV layers:



IASI has potential for detecting stratospheric WV:

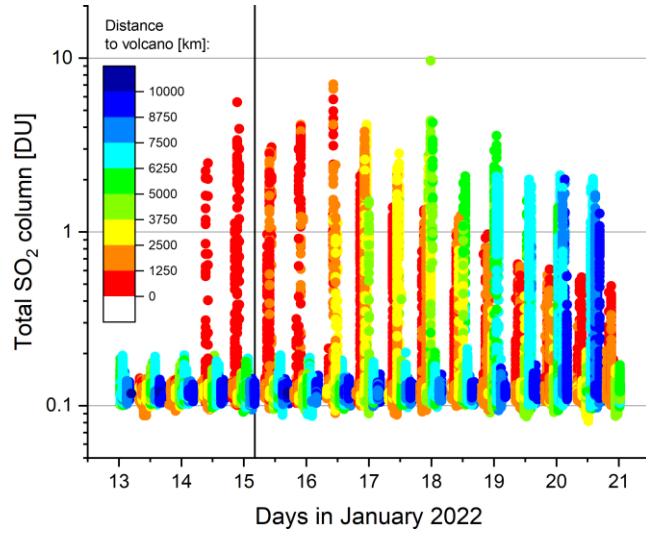
Increasing sensitivity from 108 to 5 hPa, because of increasing thermal contrast.

Retrieval setting  
(used spectral windows):

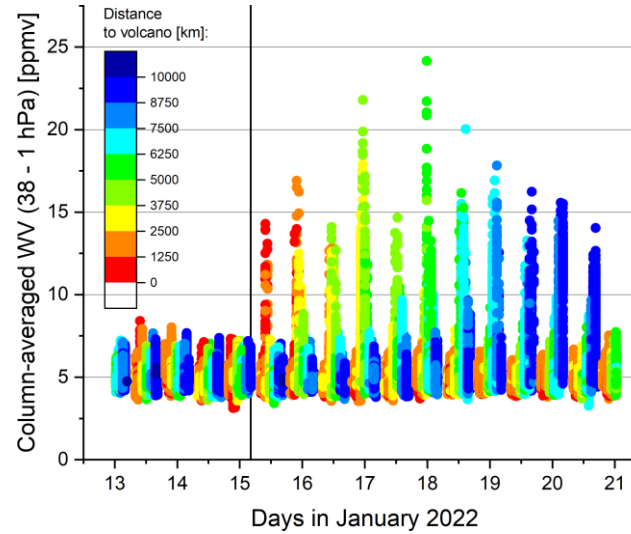


# Results: comparison of the HTHH and Raikoke volcanic eruptions

## HTHH, January 2022



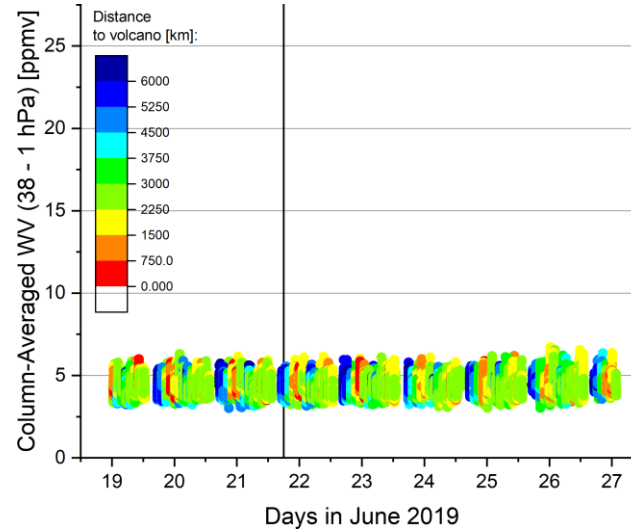
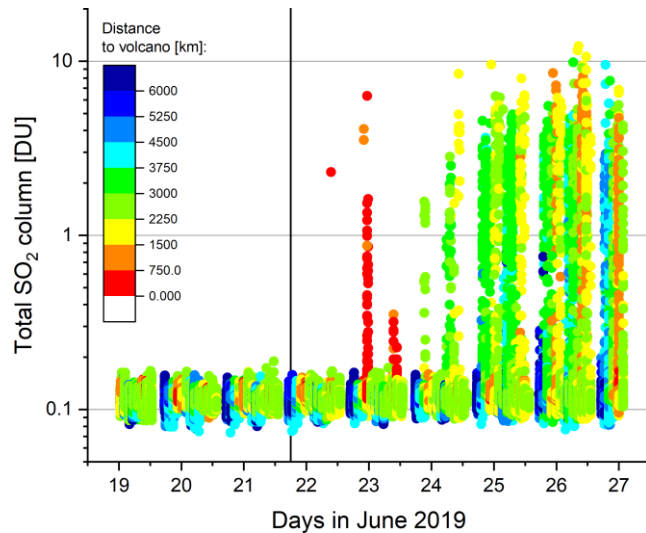
## Stratospheric Water Vapour



Data are filtered for stratospheric thermal contrast larger than 5K.

HTHH: significant SO<sub>2</sub> emission already on 220114, but explosive eruption on 220115. H<sub>2</sub>O cloud from explosive eruption, no further H<sub>2</sub>O injection at the HTHH location.

## Raikoke, June 2019



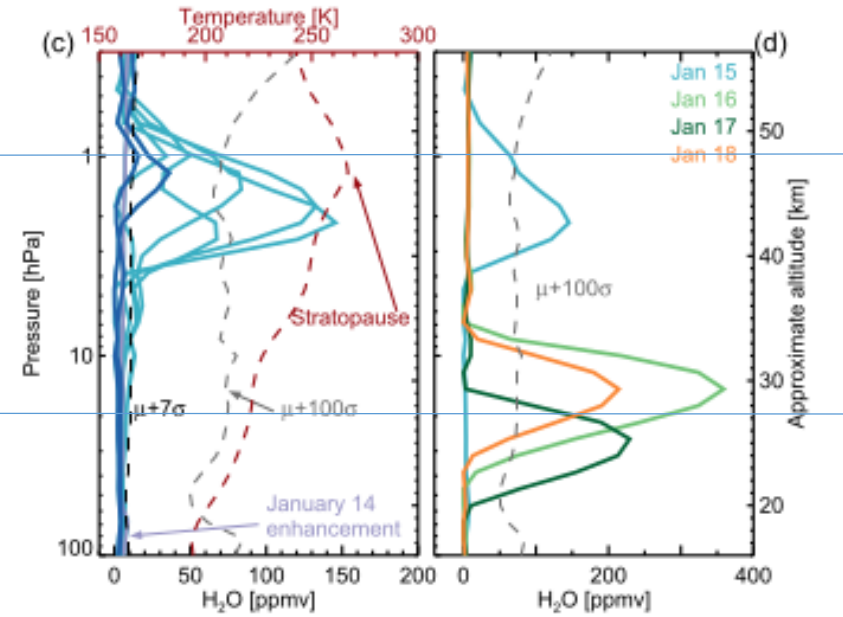
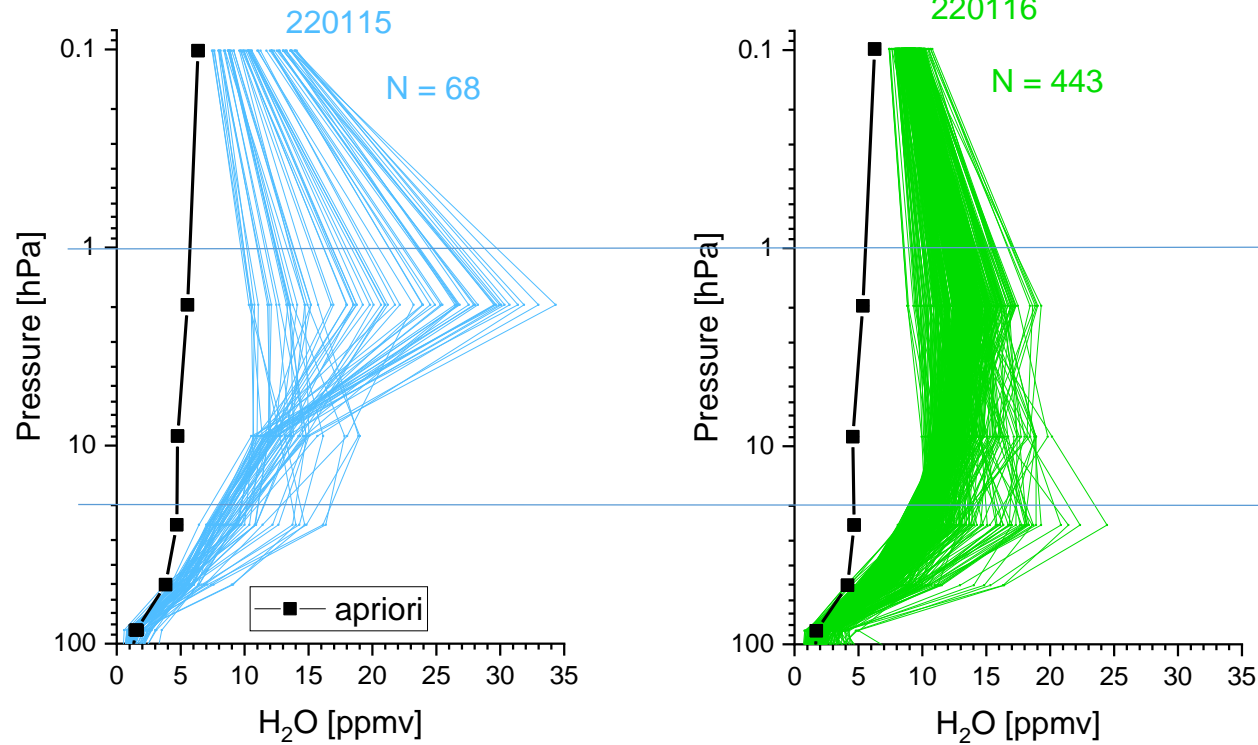
Raikoke: Similar amount of SO<sub>2</sub>, but no significant H<sub>2</sub>O injection.



# Results: comparison to MLS

Vertical anomalies detected by IASI

... and by MLS (Millán et al., 2022)

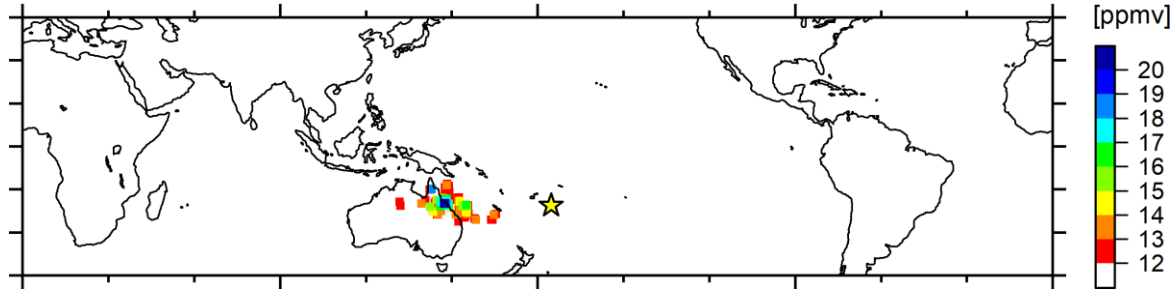


There seems to be some information on the altitude of the stratospheric WV layer.

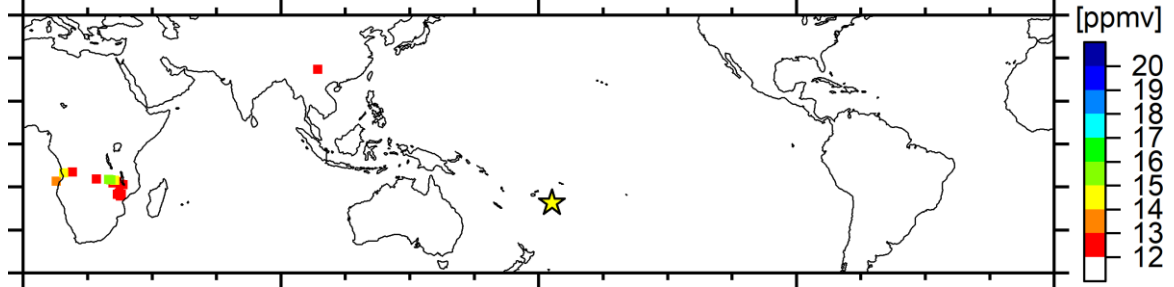
# Results: comparison to MLS

Horizontal distribution of anomalies detected by IASI

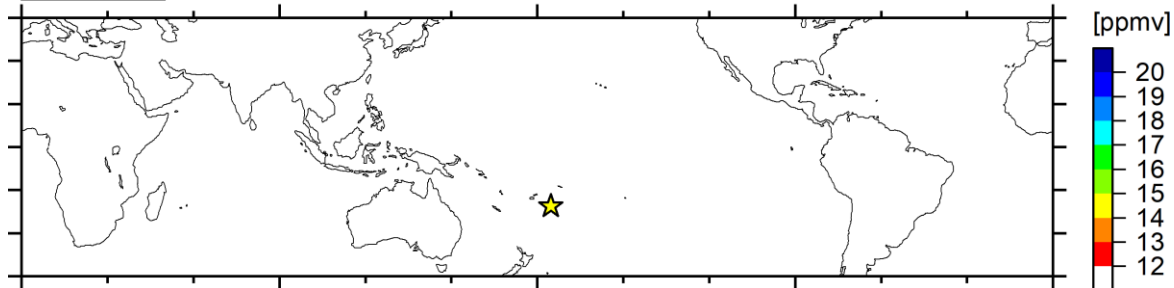
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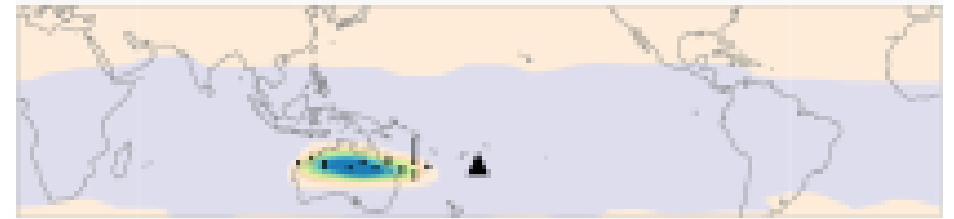
20220205



... and by MLS (Millán et al., 2022)

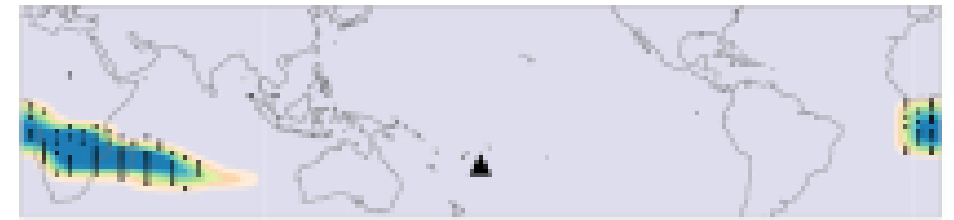
16 Jan

4.64 hPa



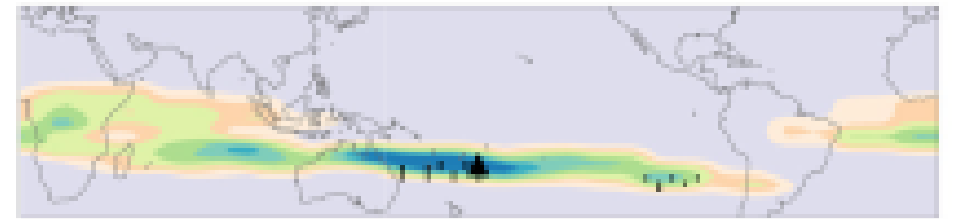
22 Jan

26.10 hPa



05 Feb

26.10 hPa



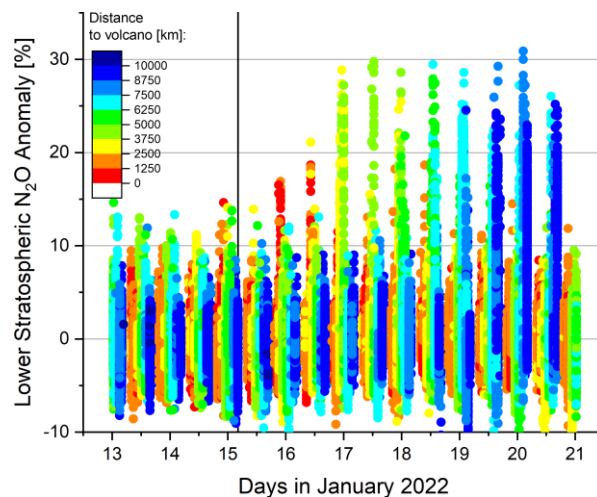
Only very strong stratospheric WV anomalies can be detected by IASI.

# Summary

- Problem: tropospheric WV and T interferences on stratospheric WV signals:
  - Constraint of RH, instead of individual constraints of WV and T.
- Thermal contrast: stratosphere is (generally) significantly warmer than the tropopause:
  - Stratospheric WV emission signal in thermal nadir radiances
- Retrieval performance check demonstrates the possibilities.
- The HTHH event is well-suited for investigating/demonstrating possibilities and limitations!

IASI/IASI-NG can (of course) not compete with dedicated limb-sounding instruments, but it can make a some contribution ...

... outlook:



Thanks for your  
attention!