

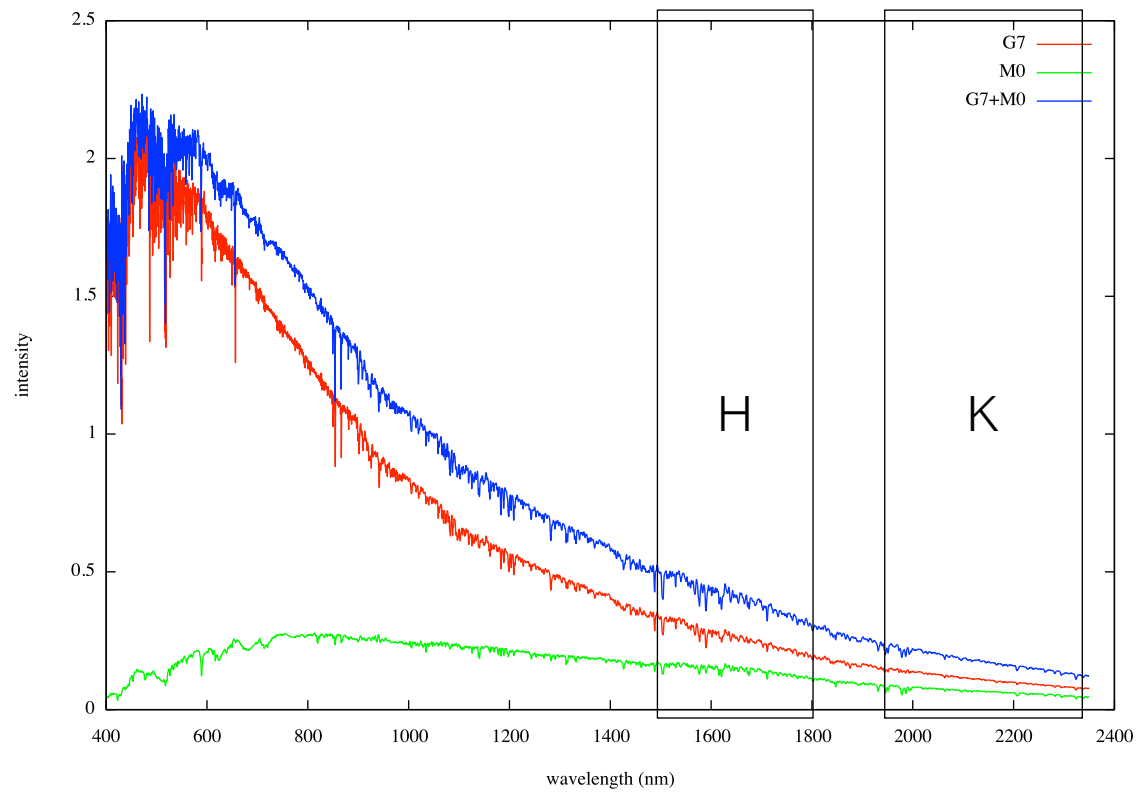
Physical Properties of the G-type Eclipsing Binaries from the Kepler Observations

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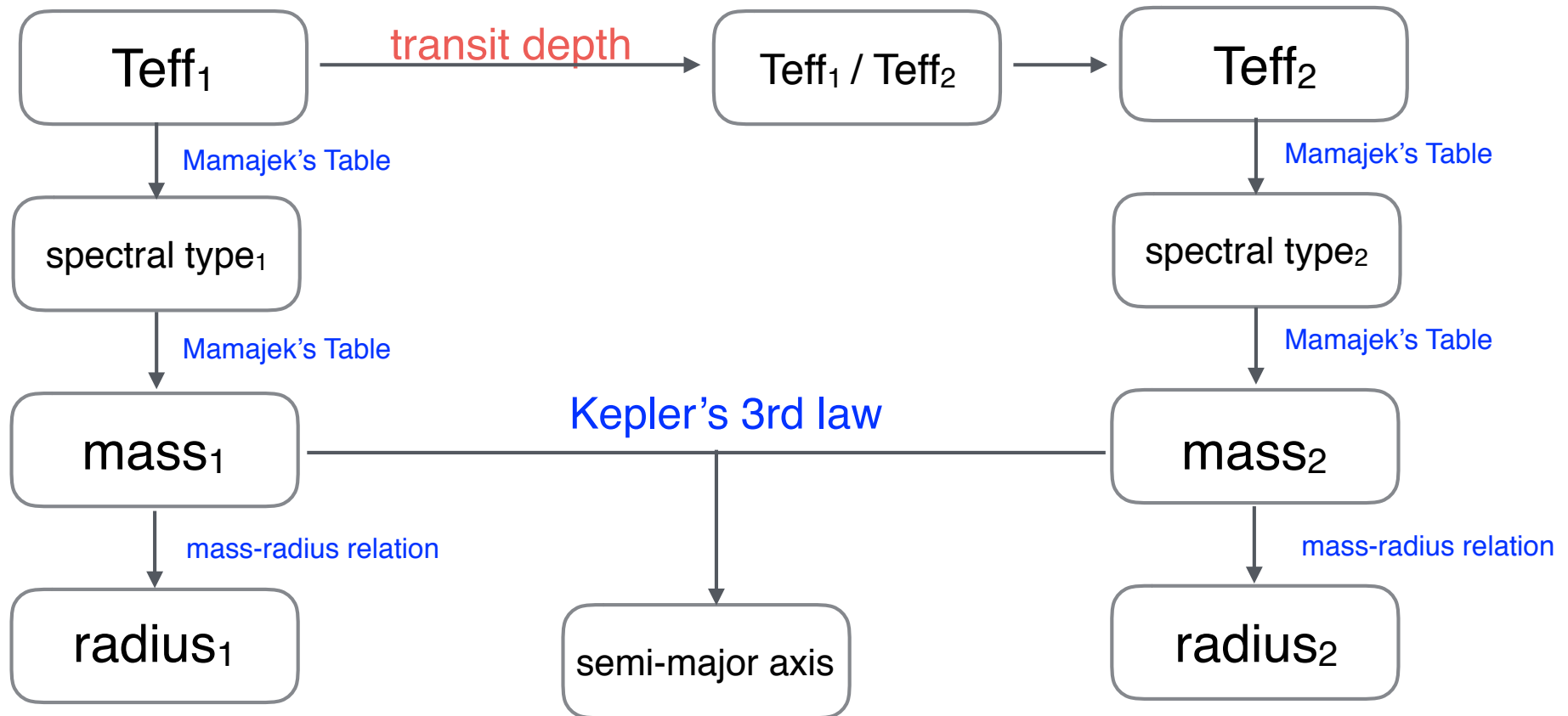
Ali Luo and Yihan Song
National Astronomical Observatories, Chinese Academy of Sciences

August 2, 2017
2nd LAMOST-Kepler Workshop, Brussels

NIR Color Table



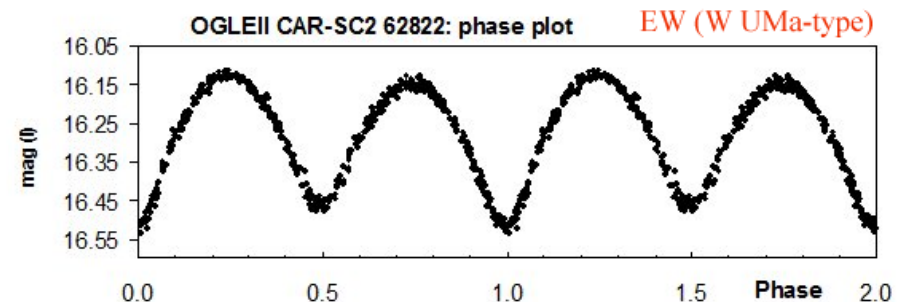
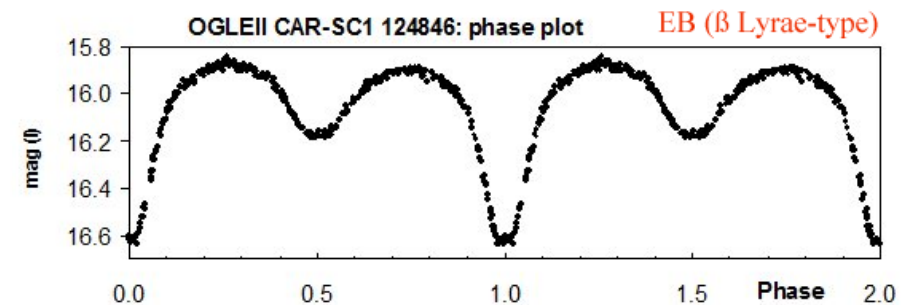
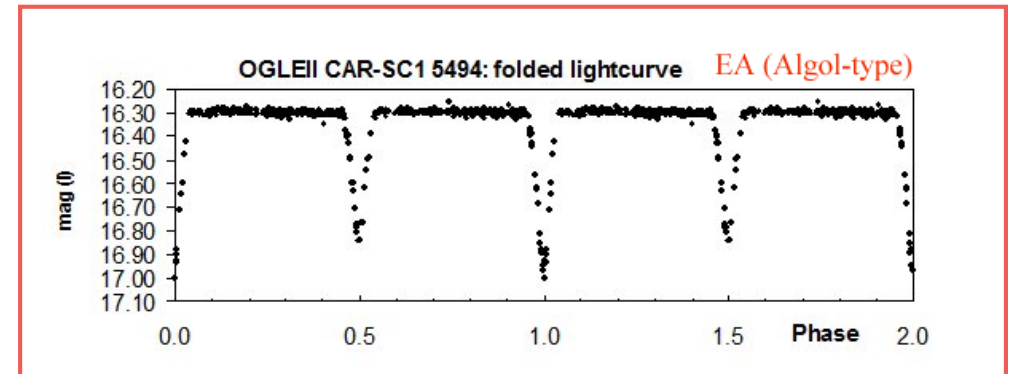
Physical Properties of Binary Components



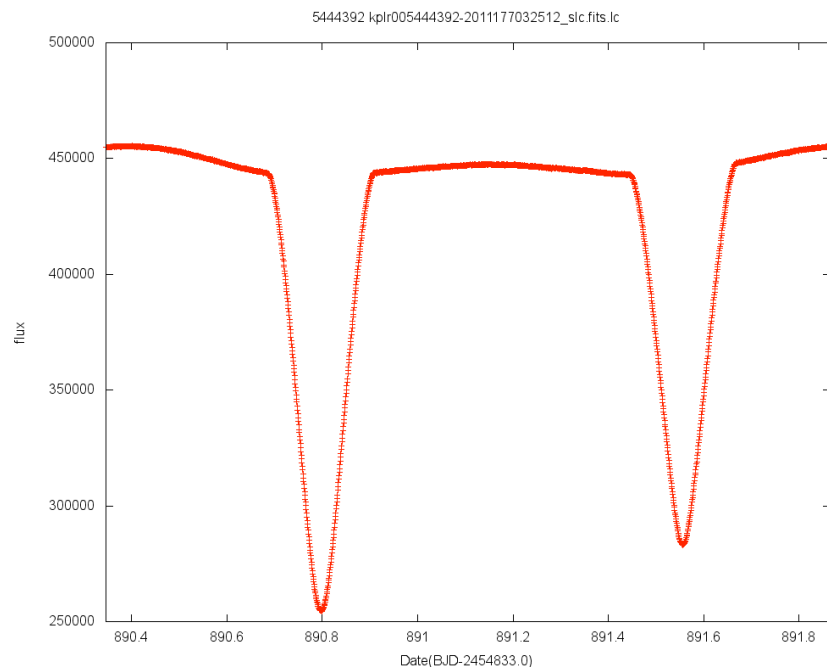
Mamajek's "Stellar Color/Teff Table" (<http://www.pas.rochester.edu/~emamajek/>)

Algol Type Eclipsing Binaries

- Teff: 5000K - 6000K (according to Huber, 2014)
- $\log(g) > 4$
- 211 Algol type EBs are selected.
- 131 have been observed by LAMOST
- 11 of them are with flare events



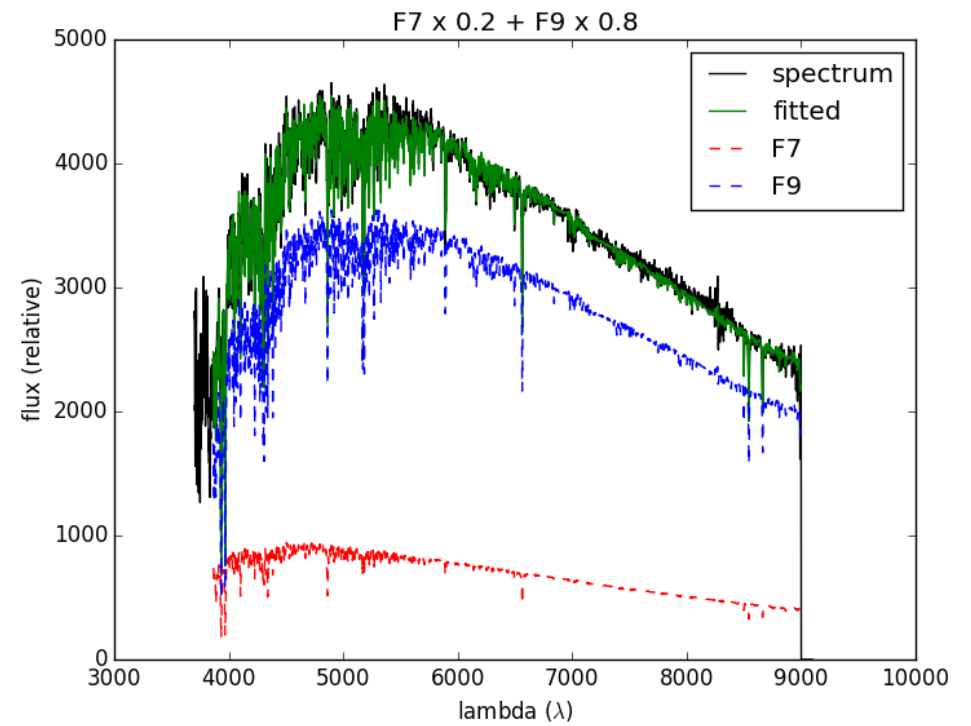
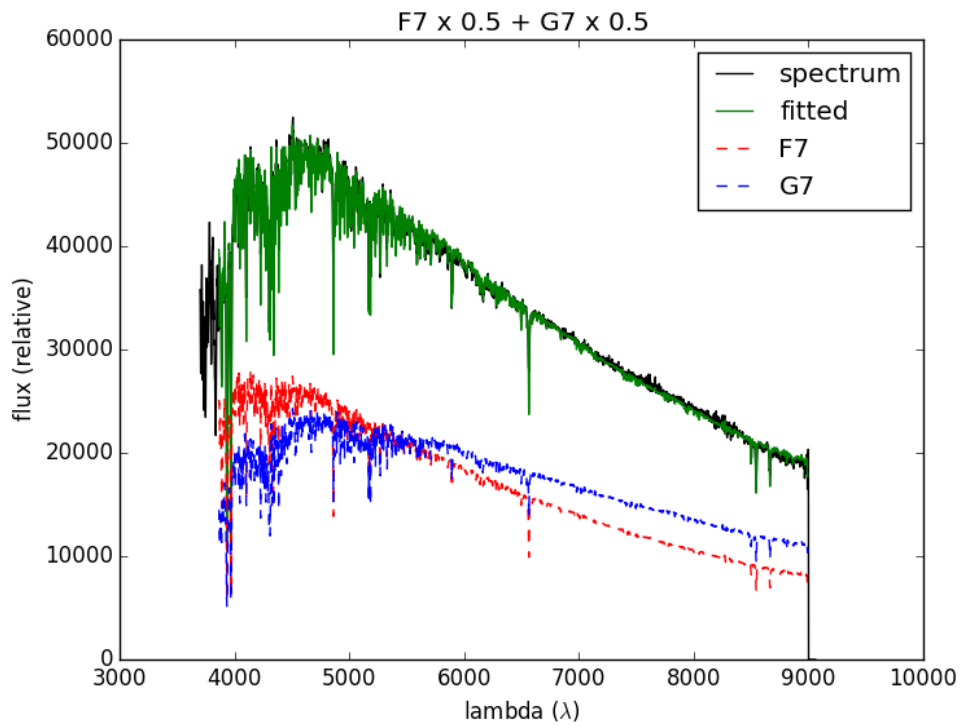
KIC 5444392



	Teff (K)	Spectral type	Mass (Msun)	Radius (Rsun)	Radius (AU)
Star1	5930	G0	1.09	1.05	0.0049
Star2	5680	G4	0.99	0.99	0.0046

Porb	T1/T2	q	R1+R2	a	flare	Tt(e)	Tt(l)
1.52	1.044	0.91	0.0095	0.028		0.17	0.20

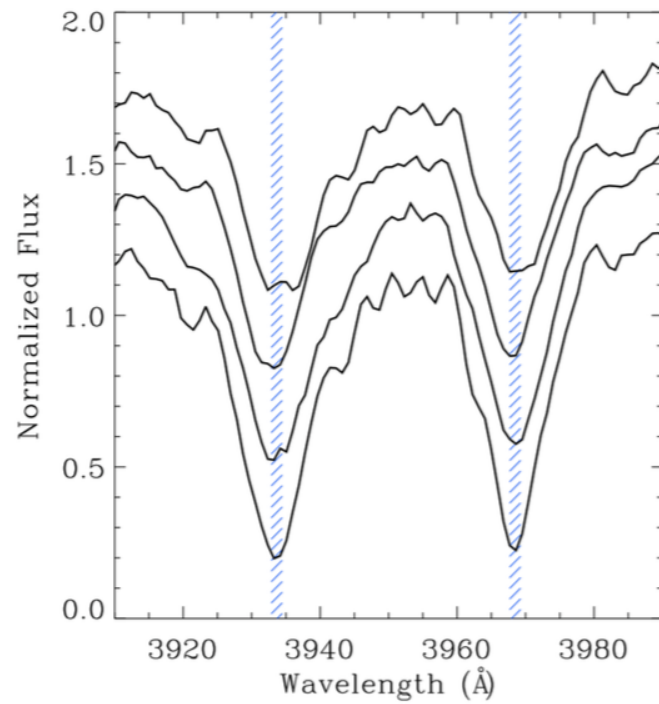
KIC 5444392 (LAMOST spectra)



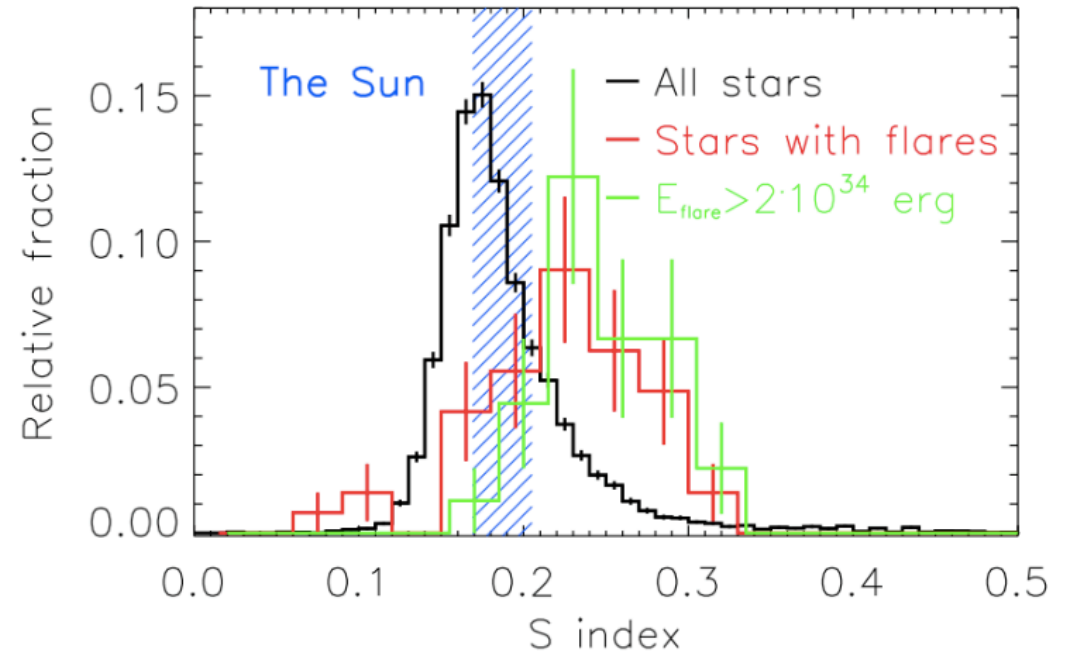
- Kepler lightcurve: 5930K (G0) + 5680K (G4)

S-index

- $S = a \cdot (H+K)/(V+R)$



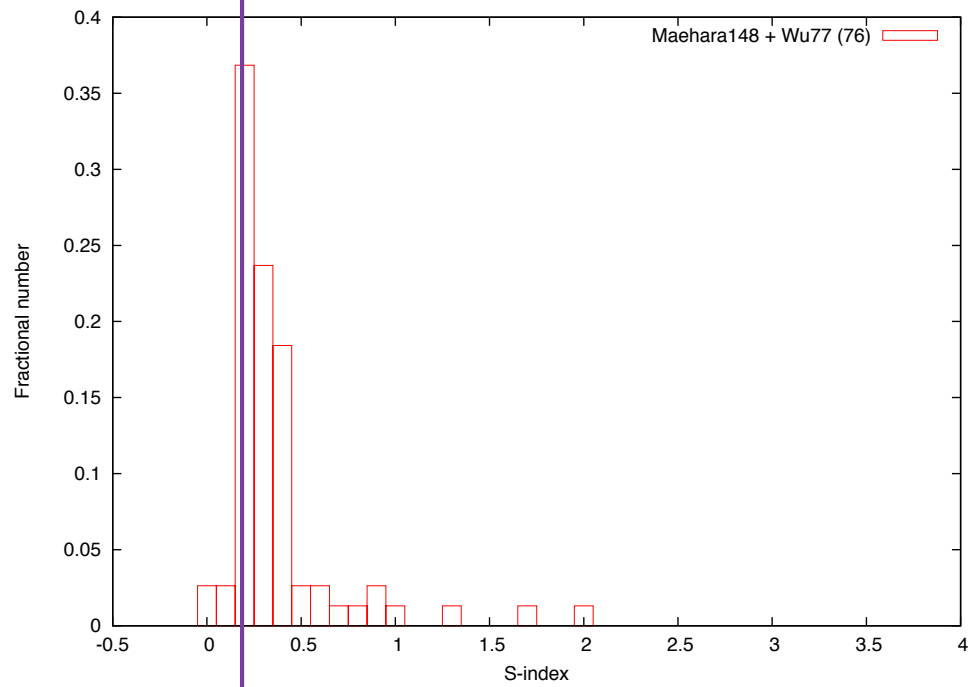
(Karoff et al., 2016)



(Karoff et al., 2016)

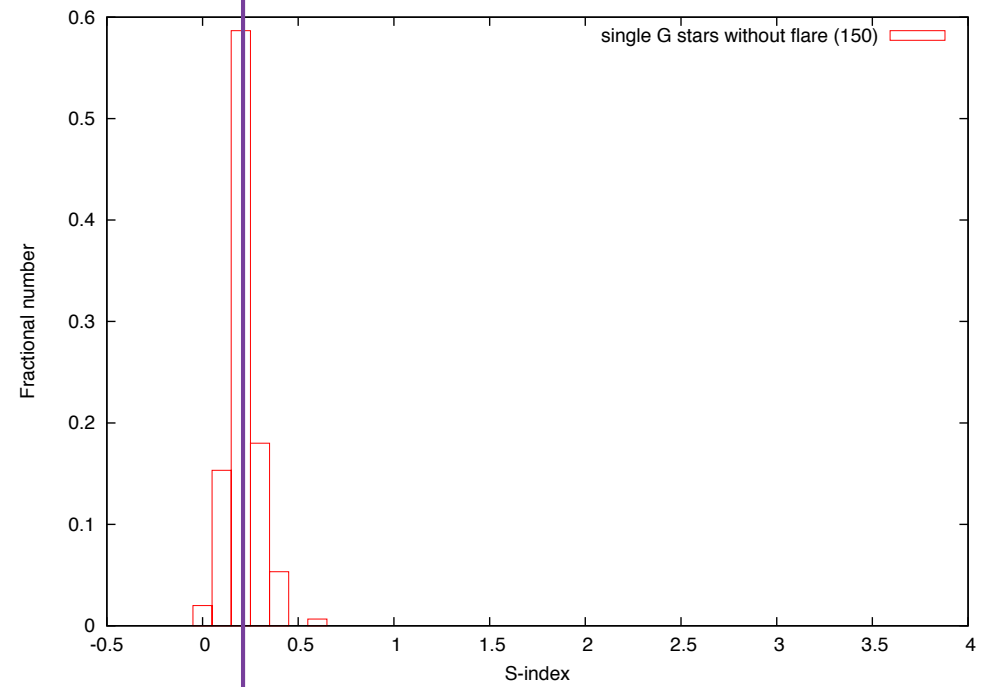
Single G-type Stars

With flares



0.2

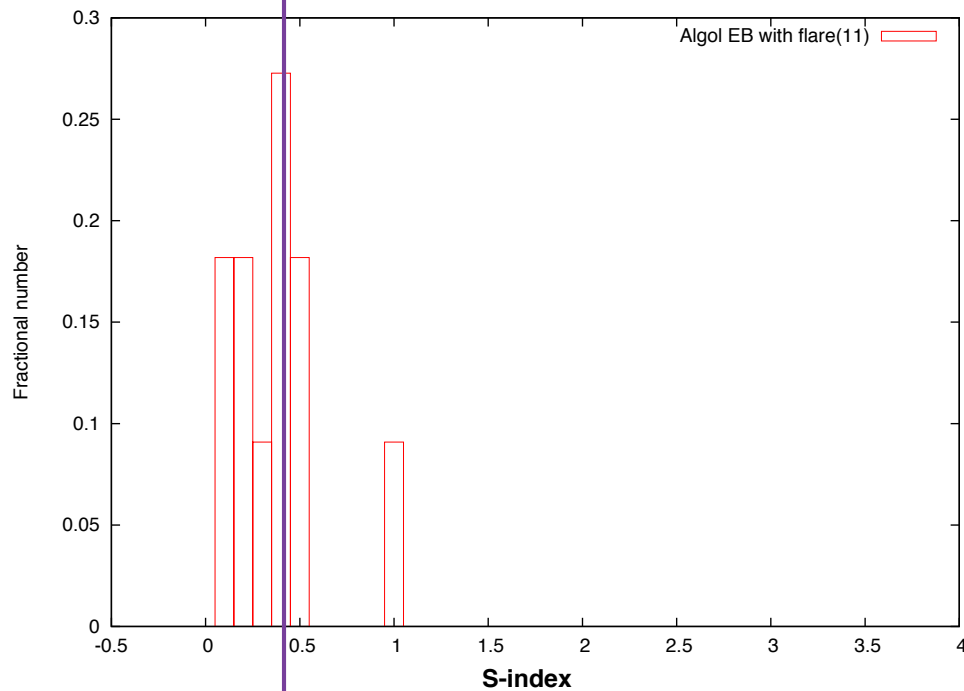
Without flare



0.2

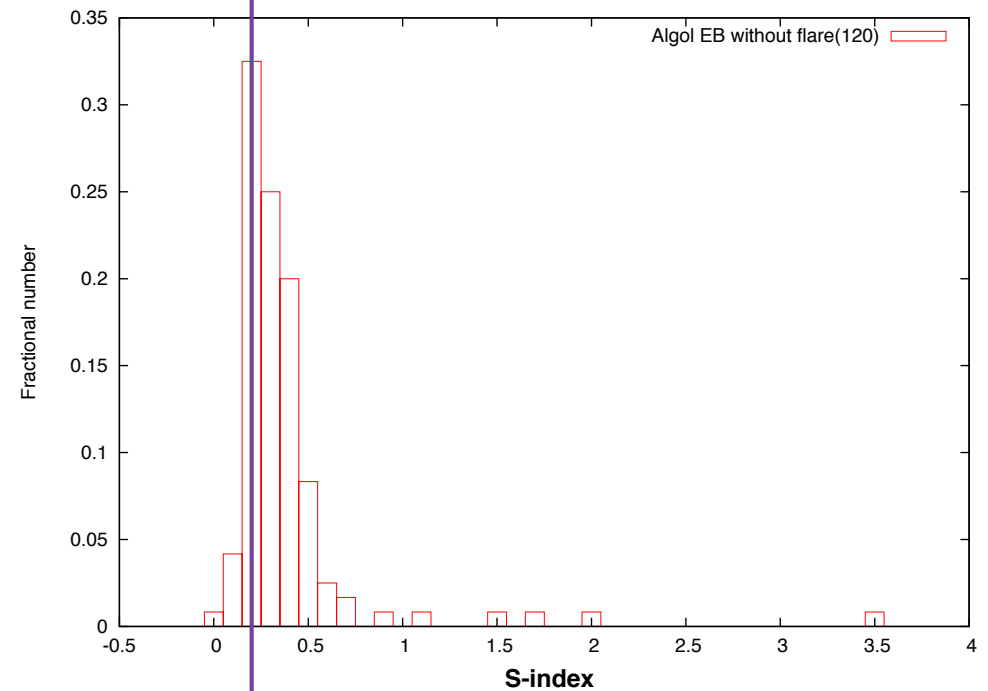
Eclipsing Binaries with G-type Primary Stars

With flares



0.4

Without flare



0.2

Thank you!

